<u>CHAPTER-III</u>

SPECIAL SPACE-LIKE CONGRUENCES ON THE STREAM LINES.

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Introduction :

In this Chapter we find the effect of vanishing of each of the curvature scalars on the space-like congruences associated with the stream line of a particle. It is argued that $K_3 = 0$ corresponds to the path of a classically gravitationally self interacting spin particle, while $K_2 = 0$ implies that the particle moves as a charged particle with radiation reaction but no external electromagnetic field. The case of $K_1 = 0$ yields a geodesic path. Section 2, portrays the three kinematical parameters of the space-like vector field P^a only after analysing the transport laws governing their definition. Explicit evaluation of physical components of shear and rotation is accomplished here. The transport laws of the spacelike congruence P^a are expressed in terms of Ricci rotation coefficients γ_{lhk} in the next section. The generalized Serret-Frenet formulae and physical components of the shear and rotation look elegant when expressed in γ_{1hk} .

Section 1 :

Special curvatures and their significance :

<u>Case 1</u> : The first curvature vanishes :

Here $K_1 = 0$, that is the magnitude of P^a vanishes. This is possible only when $\dot{u}^a = 0$. This represents the path of a free particle, in other words, it denotes trajectory of a particle

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upon which no force acts apart from the gravitational force. This path is also referred as the geodesic path.

<u>Case 2</u> : <u>The second curvature vanishes and the first</u> <u>curvature does not vanish</u> :

The equation

$$K_2 = 0, \quad K_1 \neq 0$$

implies that

 $W^{a} = 0$ by (2.18)

which means that

$$Q^a = 0$$

or $\frac{\mathbf{\dot{u}}^{a}}{K_{1}} - \frac{K_{1}}{K_{1}^{2}} \mathbf{\dot{u}}^{a} - K_{1} \mathbf{u}^{a} = 0$

or

 $\dot{u}^{a} = \frac{\dot{K}_{1}}{K_{1}} \dot{u}^{a} + K_{1}^{2} u^{a}$ (3.1)

Here \dot{u}^a is a linear combination of \dot{u}^a and u^a . Such as situation exists in the case of the path of a charged particle with radiation reaction but no external electromagnetic field. Specifically

$$m\dot{u}^{a} = \frac{e^{2}}{6\pi c^{2}} \left[\dot{u}^{a} + u^{a} \left(- \dot{u}^{b} \dot{u}_{b} \right) \right]$$
(3.2)

where m is the mass and e is the charge of the particle (Barut, 1964).

<u>Remark</u> : When $K_1 = \text{constant}$, we have

$$\hat{K}_1 = 0.$$

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$$u^{a} = K_{1}^{2} u^{a}$$
. (3.3)

And this equation represents the differential equation for a time-like circle. (Synge, 1960).

<u>Case 3</u> : <u>The third curvature vanishes and the first</u> and the second do not vanish :

Here

$$K_3 = 0, K_2 \neq 0, K_1 \neq 0.$$
 (3.4)

This is similar to the previous case, we get that \ddot{u}^a should be a linear combination of \ddot{u}^a , \dot{u}^a , u^a that is

$$\ddot{\mathbf{u}}^{a} = \lambda \ \ddot{\mathbf{u}}^{a} + \mu \ \dot{\mathbf{u}}^{a} + (\lambda \ \dot{\mathbf{u}}^{b} \dot{\mathbf{u}}_{b} - 3 \ \mathbf{u}^{b} \ddot{\mathbf{u}}_{b}) \ \mathbf{u}^{a}$$
(3.5)

where λ and μ are arbitrary.

Such a situation exists in the path of a classically gravitationally self interacting spin particle with Frankel-Weyssenhoff constraints, viz.,

$$\mathbf{\ddot{u}^{a}} + (3 \ \mathbf{\ddot{u}^{b}}\mathbf{\ddot{u}_{b}}) \ \mathbf{u^{a}} = (\frac{m^{2}}{s^{2}} - \mathbf{\dot{u}^{b}}\mathbf{\ddot{u}_{b}}) \ \mathbf{\ddot{u}^{a}} + \frac{8}{15} \ Gm$$

$$(\frac{m^{2}}{s^{2}} - \mathbf{\dot{u}^{2}}) \ (\mathbf{\ddot{u}^{b}}\mathbf{\ddot{u}_{b}} \ \mathbf{u^{a}} + \mathbf{\ddot{u}^{a}}) \qquad (3.6)$$

(Goenner <u>et al</u>. 1967)

where
$$S_a = S_{ab} u^b$$
 with the spin tensor S_{ab}
m = mass
G = Newtonian Gravitational constant.

<u>Remark</u> :

When $K_3 = 0$ and $K_1 = \text{constant}$, $K_2 = \text{constant}$

we have

$$\dot{K}_1 = 0$$
 and $\dot{K}_2 = 0$

These conditions implies that the path of the particle is the time-like helix.

Section 2 :

Shear-free, irrotational space-like congruence Pa :

(1) The three parameters of the space-like congruence P^a:

The parameters of the space-like congruence P^a , relative to the time-like congruence u^a when the signature of metric is (--+) are cited below. We note that in Chapter I, Greenberg's formulae are for the metric signature (+++-).

(i) The expansion parameter is defined by

$$\frac{\theta}{(1)} = \frac{1}{2} \left(P^{a}_{;a} - P_{a;b} u^{a} u^{b} \right)$$
 (3.7)

here the subscript (1) below Θ denotes that the parameter is for the first space-like congruence P^a .

where \perp_{ab} is the 2-dimensional projection operator defined by

$$\mathcal{L}_{ab} = \mathcal{J}_{ab} - u_a u_b + P_a P_b \qquad (3.9)$$

with properties

(iii) The rotation tensor field for P^a is characterized by

$$\begin{aligned} \omega_{ab} &= \perp_{a}^{c} \perp_{b}^{d} (P_{c;d} - P_{d;c}). \end{aligned}$$
(3.11)

These definitions are subject to the three GREENBERG'S transport laws for P^a (after due corrections for signature):

$$u^{a}_{;b} P^{b} = P^{a}_{;c} u^{c} - u^{a} P_{b;c} u^{b} u^{c} + P^{a} P_{b;c} u^{b} P^{c}$$
 (3.12)

$$Q^{a}_{;b}P^{b} = -u^{a}P_{b;c}Q^{b}u^{c} + P^{a}P_{b;c}Q^{b}P^{c}$$
 (3.13)

and

$$R^{a}_{;b} P^{b} = -u^{a} P_{b;c} R^{b} u^{c} + P^{a} P_{b;c} R^{b} P^{c}$$
. (3.14)

Since the index a ranges over () to 3, these are 12 equations (3 of which will be shown to be identities). The significance of these transport laws is that, they ensure the orthogonality of the tetrad (u^a , P^a , Q^a , R^a) during the parallel transport of the vector fields.

(2) The expansion, the shear and the rotation of the space-like congruence P^a in terms of u^a , \dot{u}^a , \ddot{u}^a :

(i) By the expressions (2.5), (3.7) becomes

$$\theta_{(1)} = \frac{1}{2} \left(\frac{\dot{u}^{a}_{;a}}{K_{1}} - \frac{K_{1;a}}{K_{1}^{2}} - K_{1} \right)$$
 (3.15)

since $\dot{u}^a u_a = K_1^2$ and $\dot{u}^a u_a = 0$, where K_1 is the first curvature of the world line.

(ii) Using (2.5) in (3.8) we have

$$\int_{(1)ab}^{0} = \perp_{a}^{c} \perp_{b}^{d} \left(\frac{\dot{u}_{c;d}}{K_{1}} - \frac{K_{1;d} \dot{U}_{c}}{K_{1}^{2}} + \frac{\dot{u}_{d;c}}{K_{1}} - \frac{K_{1;c} \dot{u}_{d}}{K_{1}^{2}} \right)$$
$$- \perp_{ab}(\theta)$$

and by using (3.9) in above expression

$$\begin{aligned}
\left(\int_{(1)}^{\infty} ab &= \frac{1}{K_{1}} \left(\dot{u}_{a;b} + \dot{u}_{b;a} - u_{a}\dot{u}_{b} - \dot{u}_{a}u_{b} - u_{a}\dot{u}_{c;b} u^{c} - u_{b}\dot{u}_{c;a}u^{c} \right) + \frac{\dot{k}_{1}}{K_{1}^{2}} \left(\dot{u}_{a}u_{b} + u_{a}\dot{u}_{b} \right) + 2K_{1}u_{a}u_{b} \\
&+ \frac{1}{K_{1}^{3}} \left(\dot{u}_{a} \dot{u}_{b;c} \dot{u}^{c} + \dot{u}_{a} \dot{u}_{c;b} \dot{u}^{c} + \dot{u}_{a}\dot{u}_{c;b} \dot{u}^{c} + \dot{u}_{a;c} \dot{u}_{c;b} \dot{u}^{c} + \dot{u}_{a;c} \dot{u}_{c;c} \dot{u}^{c} u^{d} - u_{a} \dot{u}_{b} \dot{u}_{c;d} \dot{u}^{c} \dot{u}^{d} \\
&+ \dot{u}_{a;d} \dot{u}_{b} \dot{u}^{d} + \dot{u}_{d;a} \dot{u}_{b} \dot{u}^{d} \right) \\
&+ \frac{1}{K_{1}^{5}} \dot{u}_{a}\dot{u}_{b}\dot{u}^{c}\dot{u}^{d} \left(\dot{u}_{c;d} + \dot{u}_{d;c} \right) - \int_{ab}^{0} \dot{u}_{(1)}^{0} \cdot (3.16)
\end{aligned}$$

Note : On the non-zero independent components of $\int_{(1)}^{(1)} db$:

Since
$$\int_{ab} u^{a} = 0$$
 we have
 $(1)^{ab} u^{a} = 0$ (3.17)

and therefore (1)ab is orthogonal to u", which is

Now $\perp_{ab} P^a = 0$

implies that

$$\perp_{ab} \dot{u}^a = 0$$

therefore

hence $\delta_{(1)ab}$ is orthogonal to u^a .

From (3.17) and (3.18) it follows that $(1)^{ab}$ is in the 2-plane spanned by Q^a, R^a. Consequently there are at most 2-non-zero components of $(1)^{ab}$, since N (3.19)

and

(iii) Now using (2.5) in (3.11) we have

$$\overset{(U)}{(1)}_{ab} = \perp_{a}^{c} \perp_{b}^{d} \left[\frac{1}{K_{1}} (\dot{u}_{c;d} - \dot{u}_{d;c}) + \frac{1}{K_{1}^{2}} \right]$$

$$(K_{1;c} \dot{u}_{d} - K_{1;d} \dot{u}_{c})$$

and by using (3.9) in the above expression, we get

$$\begin{aligned} & (i)_{(1)ab} = \frac{1}{K_{1}} (\dot{u}_{a;b} - \dot{u}_{a}u_{b} - u_{a}\ddot{u}_{b} + \ddot{u}_{d;a} \dot{u}_{b}u^{d} - u_{a}\dot{u}_{c;d}u^{c}) \\ & + \frac{\dot{K}_{1}}{K_{1}^{2}} (\dot{u}_{a}u_{b} - u_{a}\dot{u}_{b}) + \frac{1}{K_{1}^{3}} (\dot{u}_{a}\dot{u}^{c} (\dot{u}_{c;b} - \dot{u}_{b;c})) \\ & + \dot{u}_{a}u_{b}\dot{u}_{d;c} \dot{u}^{c}u^{d} - u_{a}\dot{u}_{b}\dot{u}_{c;d}\dot{u}^{d} + (\dot{u}_{a;d} - \dot{u}_{d;a})) \\ & \dot{u}_{b}\dot{u}^{d}) + \frac{1}{K_{1}^{5}} \dot{u}_{a}\dot{u}_{b}\dot{u}^{c}\dot{u}^{d} (\dot{u}_{c;d} - \dot{u}_{d;c}) \quad (3.20) \end{aligned}$$

<u>Note</u>: Since $\perp_{ab} u^{a} = 0$ and $\perp_{ab} P^{a} = 0$, $(\bigcup_{\substack{(1)\\ab}} u^{a} = 0$ (3.21)

and

i.e. $(\mathcal{U})_{ab}$ is orthogonal to u^a and \dot{u}^a , (1)

 $\bigcup_{(1)^{ab}} \mathbf{\dot{u}}^{a} = 0$

From (3.21) and (3.22) it follows that $(\mathcal{W}_{(1)^{ab}})$



is

in the 2-plane spanned by Q^a , R^a . There is atmost one non-zero component of (\mathcal{U}_{ab}) , since

The vorticity space-like congruence ω^a is given by

$$\omega^{a} = \frac{1}{2} \eta^{abcd} u_{b} \omega_{cd}$$
 (3.24)

(3) Physical Components of a tensor :

We define the physical components of a tensor A abcd to be the set of scalars

$$A_{\alpha\beta\gamma\delta} = e_{(\alpha)}^{(\alpha)} e_{(\beta)}^{(b)} e_{(\gamma)}^{(c)} e_{(\delta)}^{(d)} A_{abcd} \qquad (3.25)$$

where Greek indices range over 0,1,2,3 and

$$e^{a}_{(\prec)} = \{ u^{a}, P^{a}, Q^{a}, R^{a} \}$$
.

(i) Physical components of $(5)^{ab}$:

From (3.25), we write $\int_{(1)}^{0} \langle \beta \rangle = \begin{pmatrix} a \\ (a \end{pmatrix} \begin{pmatrix} b \\ (b \end{pmatrix} \end{pmatrix} \begin{pmatrix} \beta \end{pmatrix} \begin{pmatrix} \beta$

We have atmost two non-zero components of $\mathcal{O}_{(1)ab}$ and so we evaluate $\mathcal{O}_{(2)}$, viz.,

$$(1)^{22} = e^{a}_{(2)} e^{b}_{(2)} (1)^{ab}$$
$$= Q^{a} Q^{b}_{(1)ab}$$

by (3.16), above expression becomes

$$(\overset{\circ}{1})^{22} = Q^{a} Q^{b} \frac{1}{K_{1}} (\overset{\circ}{u}_{a;b} + \overset{\circ}{u}_{b;a}) - \perp_{ab} Q^{a} Q^{b} \Theta$$
(1)

since $Q^a u_a = 0$, $Q^a \dot{u}_a = 0$.

or

$$\int_{(1)^{22}} = \frac{2 \tilde{u}_{a;b}}{K_1} Q^a Q^b - g_{ab} Q^a Q^b \theta$$
(1)

i.e.
$$- \int_{(1)^{33}} = \int_{(1)^{22}} = \frac{2}{K_1} \dot{u}_{a;b} Q^a Q^b + \Theta$$
 (3.26).

Now, from (3.25)

$$(1)^{23} = Q^{a} R^{b} (1)^{ab}$$

by (3.16), we have

•

$$(1)^{23} = Q^{a} R^{b} \frac{\overset{u}{a;b} + \overset{u}{u}_{b;a}}{\overset{R}{n_{1}}} - \bot_{ab} Q^{a} R^{b} \overset{\Theta}{(1)}$$

the equation (3.9), (3.19) gives that

$$\int_{(1)^{23}}^{\infty} = + \int_{(1)^{32}}^{\infty} = \frac{1}{K_1} (\dot{u}_{a;b} + \dot{u}_{b;a}) Q^a R^b.$$
 (3.27)

We summerize these results -

<u>Note</u>: Shear free P^a is characterized by the two conditions $(5)_{22} = 0, \quad (5)_{23} = 0$, these are satisfied

when P^a is a killing vector field.

(ii) Physical components of $(1)_{ab}$:

From (3.25), we write

but we have only one non-zero independant component of $(1)^{\omega}$ ab,

i.e.
$$(\mathcal{U}_{(1)}^{23} = Q^a R^b (\mathcal{U}_{(1)}^{ab})$$

by using (3.20) in above equation, we have -

$$\bigcup_{(1)^{23}} = - \bigcup_{(1)^{32}} = \frac{1}{K_1} (\dot{u}_{a;b} - \dot{u}_{b;a}) Q^a R^b$$
 (3.28)

we summerize these results

Note : Irrotational congruence Pa can be described through

$$\dot{u}_{a;b} = \dot{u}_{b;a}$$

i.e. u_a is a harmonic congruence.

Section-3 :

Serret-Frenet formulae, transport laws and physical components in terms of Ricci Rotation Coefficients: :

The set of invariants γ_{lhk} , defined by the equations $\gamma_{lhk} = e_{(1)|a;b} e_{(h)|}^{a} e_{(k)|}^{b}$ (Eisenhart, 1960) ... (3.29)

where 1, h, k range over (0, 1, 2, 3) and where

$$e_{(0)}^{a} = u^{a}, e_{(1)}^{a} = P^{a}, e_{(2)}^{a} = Q^{a}, e_{(3)}^{a} = R^{a}$$
 (3.30)

is called as Ricci rotation coefficients (Scalars) with properties

$$\gamma_{\rm lhk} + \gamma_{\rm hlk} = 0 \tag{3.31}$$

$$\gamma_{llk} = 0 \qquad (l \text{ is not dummy}) \cdot \qquad (3.32)$$

- (2) <u>Transport laws of the space-like congruence</u> P^{a} in <u>terms of u^a, u^a, u^a :</u>
 - (i) By using (2.5) in (3.12) we have

$$u^{a}_{;b} P^{b} = \frac{u^{a}}{K_{1}} - \frac{K_{1}}{K_{1}^{2}} u^{a} - K_{1} u^{a} + \frac{1}{k_{1}^{3}} u^{a} u^{b}_{b;c} u^{c} u^{b}$$

or

$$u^{a}_{;b} P^{b} = K_{2} Q^{a} + (\frac{1}{K_{1}^{2}} \dot{u}_{b;c} \dot{u}^{c} u^{b}) P^{a}$$
 (3.33)

We now express these 4 equations in terms of $\gamma_{\rm lhk}$.

(ii) By using (2.5) in (3.13), we have

$$Q^{a};_{b} P^{b} = K_{2} u^{a} + \left(\frac{1}{K_{1}K_{2}} \dot{u}_{b;c} \dot{u}^{c} Q^{b}\right) P^{a}.$$
 (3.34)

(iii) By using (2.5) in (3.14), we get

$$R^{a}_{;b} P^{b} = \frac{1}{K_{1}^{5}K_{2}} \dot{u}^{a} \dot{u}_{b;c} \dot{u}^{c} \eta^{b lmn} u_{l} \dot{u}_{m} \ddot{u}_{n}$$

i.e., $R^{a}_{;b} P^{b} = (\frac{1}{K_{1}^{2}} \dot{u}_{b;c} \dot{u}^{c} R^{b}) P^{a}$. (3.35)

(3) GSF formulae in terms of Ricci rotation coefficients :

By contracting expressions (2.10), (2.11), (2.12) and (2.13) with u^{a} , P^{a} , Q^{a} , R^{a} we get

$$-\gamma_{010} = \gamma_{100} = \kappa_1$$
$$-\gamma_{120} = \gamma_{210} = \kappa_2$$
$$-\gamma_{230} = \gamma_{320} = \kappa_3$$

and remaining

$$\gamma_{020} = \gamma_{030} = \gamma_{130} = \gamma_{200} = \gamma_{300} = \gamma_{310} = 0 \quad (3.36)$$

(4) Transport laws in terms of Ricci rotation coefficients :

By contracting equation (3.33) with Q^a

$$r_{021} = - \frac{K_2}{2}$$
 (3.37)

and equation (3.34), contracting with u^a

$$r_{201} = \kappa_2$$

and equation (3.35) gives

The equations (3.26), (3.27) and (3.28) give that

$$2 \gamma_{122} = \int_{(1)^{22}}^{0} - \frac{\theta}{(1)} = - \int_{(1)^{33}}^{0} - \frac{\theta}{(1)}$$

but $\theta_{(1)} = \frac{1}{2} (\gamma_{122} + \gamma_{133})$

which implies that

$$\int_{(1)^{22}}^{5} = \frac{3}{2} \gamma_{122} + \frac{1}{2} \gamma_{133}$$

and

 $(1)^{23} = (1)^{32} = r_{123} + r_{132}$ and $\bigcup_{(1)^{23}} = Y_{123} - Y_{132}$. (3.38)